

# The paradoxical aspect of the Himalayan granites

JEAN-LOUIS VIGNERESSE<sup>1</sup> AND JEAN-PIERRE BURG<sup>2</sup>

<sup>1</sup>CREGU, UMR CNRS 7566 G2R, BP 23, F-54501 Vandoeuvre Cedex France  
jean-louis.vigneresse@g2r.uhp-nancy.fr

<sup>2</sup>Geologisches Institut, ETH-Zentrum, Sonneggstrasse 5, CH-8006 Zurich, Switzerland  
jpb@erdw.ethz.ch

**Abstract:** The Miocene leucogranites at the top of the crystalline High Himalaya are commonly cited as reference examples of collision-related granites. However, they are much smaller than the Hercynian collision-related granites. Additional comparison with magmatic arcs and cordilleran-type batholiths emphasises the low rate of magma production for the Himalayan granites. We review and summarise data on the condition of segregation, ascent and emplacement of leucogranitic magmas in the High Himalaya. The plutons are small and mostly concordant with the country rocks. Thermal data indicate that they were emplaced within the ductile crust. Magma ascent and segregation has been strongly controlled by the extrusion of the High Himalaya Crystallines. Strain during extrusion provided the magma its internal high anisotropy, while aggregation of successive pulses of magma coming from chemically different sources caused chemical heterogeneity. The source region had a slow melting rate (< 13 %) at rather low temperature (below 780 °C).

**Keywords:** granite emplacement, Himalaya, leucogranites, ductile deformation

## Introduction

Plate convergence leads to abundant magmatism that contributes to the formation of new crust and crustal recycling. A quantitative estimate of the rate of granite production under various tectonic settings helps understanding crustal evolution. Three main types of magmatism are associated with plate convergence.

- 1) Island arcs are generated over a long, eventually several hundreds millions years duration of subduction. An average rate of crustal creation can be estimated from the addition rates determined on short-lived (Marianas and Izu-Bonin) and long-lived arcs (Japan, New Zealand, New Guinea). The average is 20-40 km<sup>3</sup>/km/Ma (Fig. 1; Reymer & Schubert, 1984).
- 2) Continental arc magmas (Andean-type Pitcher, 1987) form large batholiths of imbricated plutons (for instance in the southern Sierra Nevada, Saint Blanquet et al., 2001), which represent large volumes of new material added to the crust. The cordilleran and coastal batholiths of the east Pacific extend from Alaska to Antarctica. Their bulk age spans from early Jurassic to present, but the period of intense magmatic activity is Cretaceous to early Tertiary (Vaughan, 1995; Ducea, 2001). The corresponding rate of magmatic production was about 200 km<sup>3</sup>/km/Ma for the Peruvian Coastal and the Canadian Coastal segments, about one order of magnitude more than island arcs (Fig. 1).
- 3) When continental plates collide, new magmas have compositions ranging from granodiorite to leucogranite (Hercynian type granites of Pitcher 1987). Rough estimates of the production rate of

the European Hercynian granites yield a medium figure within the trend of magmatic arcs (Fig. 1).

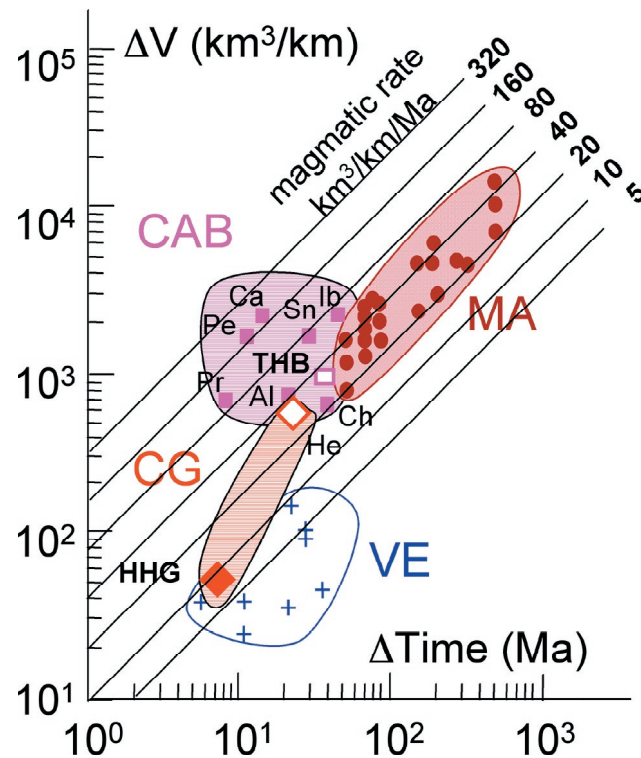
In contrast to the previous examples, the High Himalaya leucogranites have small volume and low magmatic production rate that suggest a specific generation mode. Previous workers considered either the cause of melting and/or the mode of emplacement of individual plutons. We intend to replace the Himalayan granites within a general context by specifying their characteristics and compare them to other collision-related granites, in particular the Hercynian ones. In this comparison, we particularly use the thermo-mechanical conditions of the crust, the shapes of the granitic bodies and the magmatic production rates. We term batholith huge granitic bodies composed of smaller plutons in which individual batches of magma are observed.

## Thermo-mechanical conditions of granite generation

We rapidly examine the generation conditions of granites related to continental collision. They are commonly peraluminous, derived from crustal melts, with few or no participation of mantle-derived magmas (Sylvester, 1998). We successively discuss the melting, ascent and emplacement conditions.

### *Melting conditions*

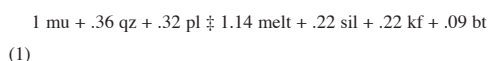
Crustal melting has long been attributed to fluid enhanced melting (Didier & Lameyre, 1969; Lameyre, 1973; Holtz & Johannes, 1994). However, dehydration melting of hydrous minerals also generates peraluminous melts (Thompson, 1982). Experiments on powder with



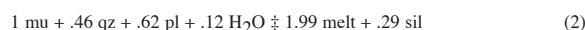
**Figure 1.** Volume of magma and crust produced in magmatic arc. Crustal production in magmatic arcs (MA) from Marianas to Izu Bonin is redrawn from Reymer & Schubert (1984). Cretaceous calc-alkaline batholiths (CAB) from the eastern Pacific are indicated (Al: Alaska, Ca Canada, Ib Idaho-Boulder, Sn Sierra Nevada, Pr Peninsula Ranges, Pe Peru coastal batholith, Ch Chilean batholith). Volcanic eruptions (VE) are also indicated. Collision granites (CG) are shown for the Hercynian (He) and Himalayas (HHG). The Trans Himalayan Batholith (THB) plots within the field of the CAB.

adequate composition (Patiño Douce & Johnston, 1991) and by re-melting peraluminous granites (Scailliet et al., 1995b) provide a frame allowing to estimate the amount of melt produced from a given source composition under given temperature and pressure conditions.

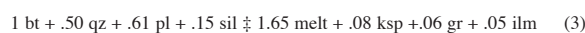
Crustal fertility can be estimated from the global composition of the source material using the equations of muscovite and biotite dehydration (Thompson, 1996; Vigneresse et al., 2001). The muscovite reaction :



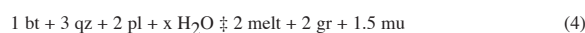
starts at about 720°C at 600 MPa (Patiño Douce & Harris, 1998) and is completed when temperatures reach about 780°C (Patiño Douce & Harris, 1998). In presence of water the reaction writes:



A unique equation for biotite dehydration melting is inadequate, though the following equation is often quoted from Patiño Douce and Harris (1998):



In presence of water the reaction becomes:



Biotite breaks down at 800-900°C with nearly no dependence on pressure. However, the amount of melt ranges from 0.28 to 2.44 % in volume depending on the initial composition of plagioclase (Patiño Douce & Harris, 1998; Harris et al., Gardien et al. 1995; Gardien et al., 2000).

We consider consequent melt development (> 15 %) to take place at 750°C for muscovite and at 850°C for biotite. We report these two significant temperatures in the thermo-mechanical profile of the crust (Fig. 2).

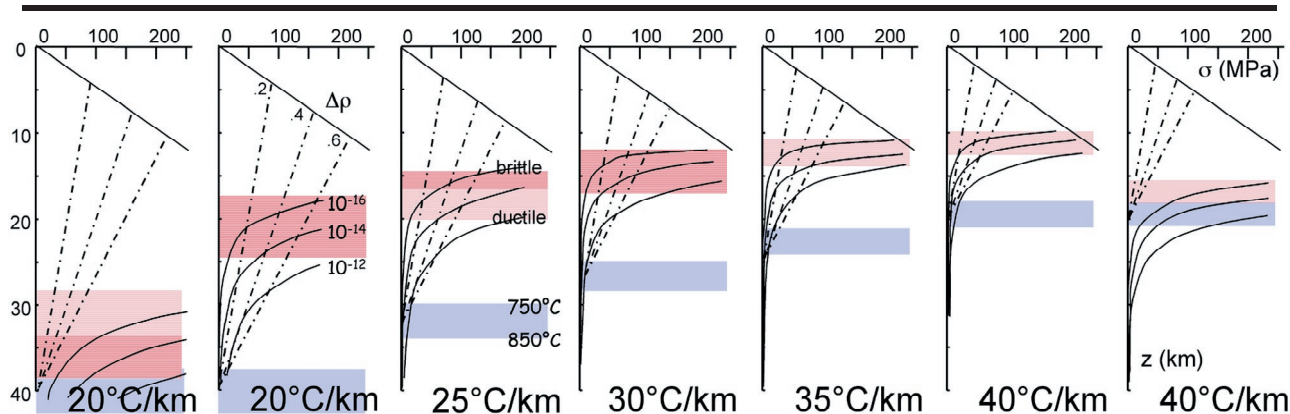
#### Ascent conditions

Melt is naturally buoyant. A column of magma formed at depth  $z_0$  develops a hydrostatic force ( $F_H$ ) that reflects the lesser density of the magma ( $\rho_m$ ) compared to the host rock ( $\rho_c$ ). Given the gravity ( $g$ ) the hydrostatic force is:

$$F_H = (z_0 - z) (\rho_c - \rho_m) g \quad (5)$$

This formulation assumes that the magma column is continuous, which is an ideal hypothesis because magma is rarely connected to its source. The ideal hydrostatic force increases linearly with the height of the magma column. We consider density contrasts between  $-0.2$  and  $-0.6 \text{ g/cm}^3$  to cover a wide range of melt / country rock densities (peraluminous magma =  $2.3 \text{ g/cm}^3$ ; crustal rocks =  $2.5 \text{ g/cm}^3$ ; granulites =  $2.9 \text{ g/cm}^3$ ; Fig. 2).

From source to emplacement depth, the magma presumably travels from the ductile lower crust to the brittle upper crust. Its ascending force must thus be compared to the stress level necessary to fracture the crust. The brittle



**Figure 2.** Bulk rheology of the crust, using a power law rheology for amphibolites (Wilks & Carter, 1990) in case of geothermal gradient varying from 20 to 40 °C/km. Extensional brittle curve only is shown. Curves of viscous rheology under different strain rate ( $10^{-12}$ ,  $10^{-14}$  and  $10^{-16}$  s $^{-1}$ ) are also indicated. It brackets the brittle/ductile transition (stippled in grey). Melting temperatures (in grey) are bracketed by the temperatures of muscovite (750 °C) and biotite (800°C) reactions. The buoyancy forces of a magma column originating at 800°C km in depth is represented using three values of density contrast (-0.2, -0.4 and -0.6 g/cm $^3$  respectively). Experimental parameters of the ductile, power-law behaviour are the activation energy ( $Q = 244$  kJ/mole), the pre-exponential coefficient ( $\log A = -9.11$  (MPa/s) $^{1/n}$ ), the power coefficient ( $n = 3.7$ ). Temperature ( $T$ ) is linked to the system by the perfect gas constant ( $R = 8.31$  J/mole/K). The geothermal gradients range from 20 to 40 °C/km.

crust respects the Griffith fracture criterium (Byerlee, 1978), with a friction coefficient  $\mu = 0.65$  to  $0.80$  (Marone, 1998). We adopt  $0.66$ . The differential stress ( $\sigma_1 - \sigma_3$ ) depends on  $\mu$  and the tectonic setting (Angelier, 1984; Vigneresse et al., 1999), which is reflected by three lines with decreasing slope from extension through strike slip to compression (Fig. 2).

To model the behaviour of the ductile crust, we adopt the power law

$$\dot{\epsilon} = \sigma_n A \exp(-Q/RT) \quad (6)$$

We take the experimentally determined parameters for amphibolites (Wilks & Carter, 1990). We bracket strain rates  $10^{-16} < \dot{\epsilon} < 10^{-12}$  s $^{-1}$  according to estimates that rule tectonic deformation (Pfiffner & Ramsay, 1982) and we compiled the rheological curves for a cold (geotherm of 20°C/km) and a hot crust (40°C/km). In figure 2, we included the two end-members values for a cold and hot crust with a rheology of wet Black Hills Quartzite and Maryland Diabase, respectively, which were used to model Himalayan tectonics (Beaumont et al., 2001).

### Emplacement conditions

Combining rheological curves corresponding to conditions discussed above, we computed the differential stress required to fracture the crust. The depth of the brittle-ductile transition occurs on a wide range of temperatures and depends on the mineralogy of, and strain rate imposed on the rocks.

### Himalayan granites

Collision of the Indian continent with Asia started at about 60-55 Ma (Besse & Courtillot, 1988; Beck et al., 1995; Patzelt et al., 1997). Four granitic belts have been recognised in the suture zone. They are, from South to North (Debon et al., 1986): the Lower Himalaya Granites

(LHG), the High Himalaya granites (HHG), the North Himalaya Granites (NHG) and the Trans-Himalaya Batholith (THB) (Fig. 3).

The southernmost LHG are dated from the late Paleozoic (Debon et al., 1986). It includes six groups of felsic magmatic rocks (Le Fort & Rai, 1999) that armour the Himalayan basement and do not belong to the collisional history.

The THB comprises several hundred kilometres long and a few tens kilometres wide units (Debon et al., 1986). Magmas range from subalkaline to meta-aluminous in composition. Located north of the ophiolite-bearing suture zone, their chemistry and age relates the THB to the northward subduction of the Indian oceanic plate below Asia. Tonalitic magma occurred more than 100 Myr ago but the major intrusion phase seems to be 54-40 Ma old (Pettersson & Windley, 1985; Debon et al., 1986; Schärer et al., 1990).

The Miocene Himalayan granites form the two other belts (Le Fort, 1981; Harrison et al., 1997). The Himalayas are the fold-and-thrust wedge developed within the Indian continent. Large-scale thrust imbrication absorbed an important part of intracontinental shortening. Seismic profiles show that the 70-80 km thick crust under the Himalayas is due to underthrusting of the Indian plate (Hirn et al., 1984; Nelson et al., 1993; Alsdorf et al., 1998).

The HHG constitute a discontinuous chain of small plutons intruded into the top levels of the High Himalaya Crystalline (HHC; Fig. 1; Le Fort, 1981; Searle, 1999). Emplacement ages range between 24-17 Ma (Scaillet et al., 1995a; Inger, 1994; Edward & Harrison, 1997; Searle et al., 1997). The chemical composition of these leucogranites and the incompletely assimilated crustal xenoliths they contain suggest a low degree of partial melting.

About 100 km to the north of the Himalayas, the NHG belt is composed of about 16 plutons (Debon et al., 1986; Harrison et al., 1997) intruded within the sedimentary series of the pre-collision continental shelf of India at 18-

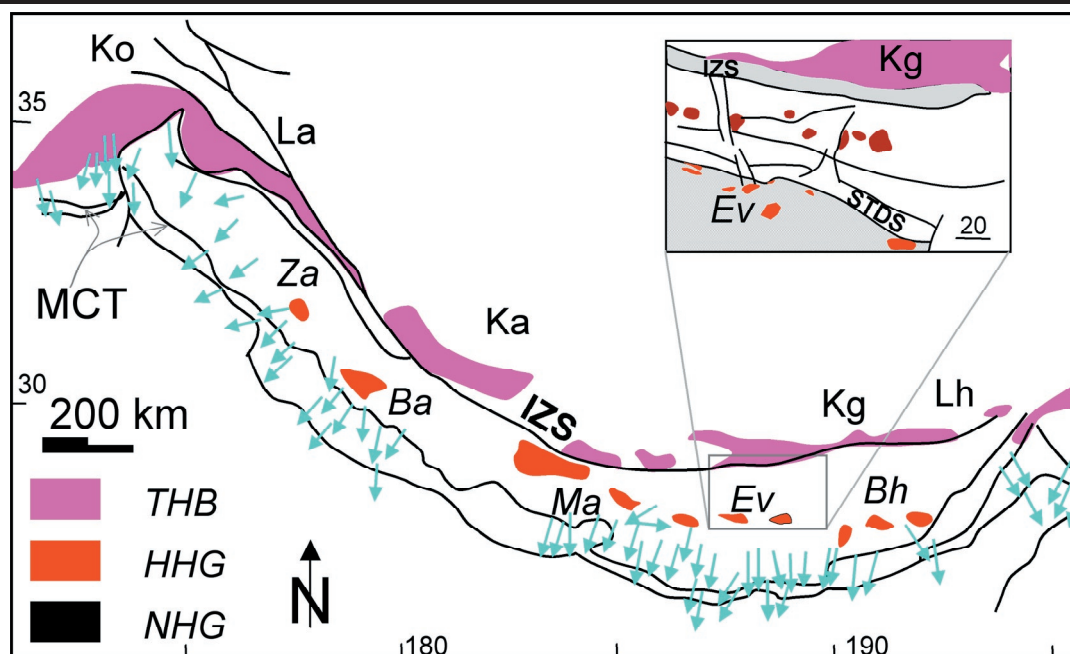


Fig. 3. Schematic map of Northern India and Nepal, showing the major structural features, including the Main Central Thrust (MCT), the South Tibetan Detachment System (STDS) and Indus Zangpo Suture (IZS), and major plutons quoted in the paper. The THB is represented in Grey by the Kohistan (Ko), Ladakh (La), Kailas (Ka), Kangdese (Kg) and Lhasa (Lh). Names of major HHG are in italic and show the Zaskar (Za), Badrinath-Gangotri (Ba), Manaslu (Ma), Everest (Ev) and Bhutan (Bh) massifs. The map for the NHG, shown in black, is detailed in the inset.

9 Ma (Schärer et al., 1986; Deniel et al., 1987; Harrison et al., 1999).

Both the NHG and the HHG have peraluminous compositions, with muscovite dominating over biotite. They typically are crust-derived magmas with few evidences of mantle-derived components.

#### Shapes of the HHG

The HHG is a 2 km to more than 10 km thick metamorphic pile interpreted as a wedge extruded southward between the bottom Main Central Thrust (MCT) and the top South Tibetan Detachment System (STDS; e.g. Hodges, 2000). The structurally lowest Formation I consists of biotite-garnet-kyanite micaschists getting migmatitic upward. The overlying and discontinuous Formation II mostly includes mid amphibolite facies carbonate rocks. Above, the Formation III is composed of sillimanite-kyanite, K-feldspar and sillimanite-cordierite-garnet metapelites screening augen orthogneiss (Le Fort, 1975; Guillot et al., 1995). The top Formation III comprises black carbonaceous and calcareous gneiss.

Sections across HHG display 1.5 to 2 km thicknesses (Searle et al., 1993; Lombardo et al., 1993; Scaillet et al., 1995a). The plutons have rather flat bases concordant with the country rocks (Fig. 3); the roofs are bounded by the normal STDS. Maps show that HHG have asymmetric shapes, reflecting shear truncation (Searle et al., 1993; Weinberg & Searle, 1999). They have 30-35 km long axes parallel to the regional stretch. Their lateral extension is more difficult to establish but the Manaslu, the largest pluton, is about 30 km long. HHG thus look like oblate ellipsoids. Calculated sizes range from 3000 km<sup>3</sup> (Manaslu) to about 150 km<sup>3</sup> (Gangotri; Scaillet et al., 1995a).

#### Thermal conditions

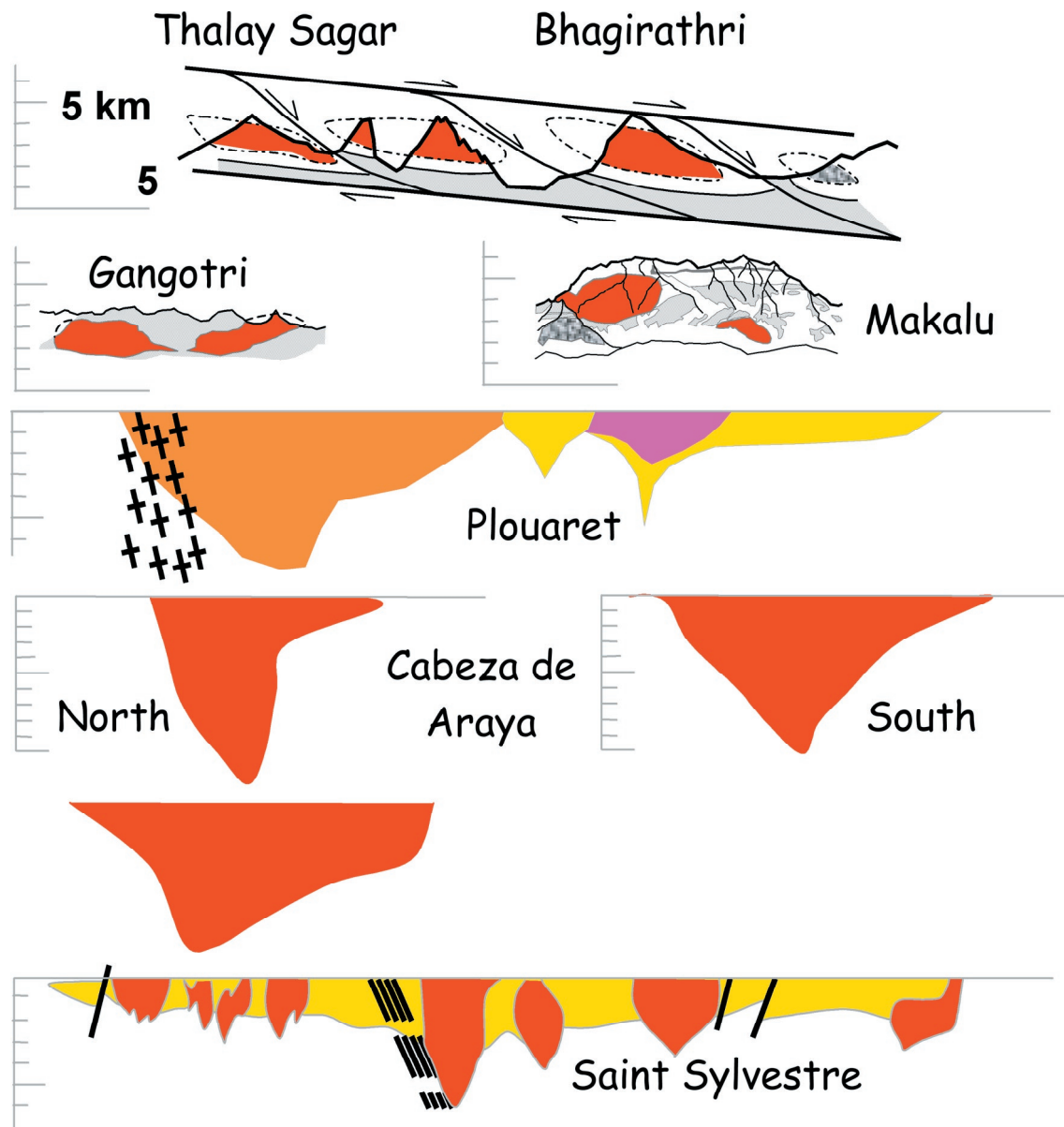
The low biotite content of HHG indicates that melt formed below the breakdown of biotite, i.e. < 800°C, presumably at 780°C (Patiño Douce & Harris, 1998). Metamorphic temperatures at the base of Formation I range 600 to 750°C with pressures around 530 MPa (Hubbard, 1989). Temperatures decrease upward to about 550°C at the top of Formation III but the pressure gradient is near lithostatic (Pêcher, 1989; Hubbard, 1989). Metamorphic aureoles yielded temperatures of 530 to 570°C (Manaslu), 440 to 575°C (Mugu) and 475-636 °C (Kula-Kangri; Guillot et al., 1995b). For the same plutons, pressure conditions are 230-400 MPa, 260-400 MPa and 290-440 MPa, respectively. Given the present thickness perpendicular to the main foliation, the geothermal gradient was about 25 °C/km. The equation of heat conduction (Carslaw & Jaeger, 1959; Furlong et al., 1991) shows that contacts reach quickly the mid temperature between those of magma and host rocks. To match estimates from aureoles, a magma temperature of 770-800 °C (Patiño Douce & Beard, 1995) implies regional temperatures of 330-400 °C (Fig. 3). These estimates are consistent with solid state and late magmatic deformation in the HHG (Burg et al., 1984), which begins around 350°C for quartz-rich rocks (Kirby & Kronenberg, 1987).

Using the geothermal gradient of 25 °C/km, melting temperatures (c. 780°C) are at 31 km depth, which fits the 775-820 MPa pressure conditions for muscovite breakdown, c. 15 km deeper than the present level of HHG.

#### Mechanical conditions

Magmatic structures in HHG are concordant with structures in the surrounding rocks (Guillot et al., 1993).





**Figure 4.** Cross section through the High Himalayan granites redrawn from photos for the Bhagirathri and Thalay Sagar (Searle et al., 1993) Makalu (Lombardo et al., 1993) and Gangotri massif (Scaillet et al., 1995a). In comparison, several cross sections, at the same scale, of Hercynian granites in Brittany (Plouaret, Pontivy), Massif Central (Saint Sylvestre) and Spain (Baseza de Araya) are provided from gravity data inversion (Améglio et al., 1997). A similar true scale (vertical = horizontal) is applied to all massifs

The general trend in the 2900 km long HHC is a N-S to N020 stretching lineation plunging about 20-30°N consistent with a bulk southward shear (Brunel, 1986, Fig.4). The magmatic fabrics are consistent from the West, in Ladakh (Weinberg, 1997) to the East, through the Zaskar (Dèzes et al., 1999), the Gangotri (Scaillet et al., 1990), the Manaslu (Pêcher et al., 1991; Guillot et al., 1993) the Chokkang arm (Pêcher et al., 1991), the Xixapanga (Searle et al., 1997), the Shisha Pangma (Searle et al., 1997), the Khumbu Himalaya (Weinberg & Searle, 1999), the Everest-Makalu (Rochette et al., 1994) and the Khula Kangri (Edwards & Harrison, 1997; Fig. 4).

Considering the 25°C/km geothermal gradient and the bulk amphibolitic composition of Formation I, the brittle/ductile transition was at 15-20 km depth (Fig. 2). For a density contrast of -0.3 (granitic magma = 2.35 g/cm<sup>3</sup> in 2.65 g/cm<sup>3</sup> Formation I) the hydrostatic magma force was stronger than the yield stress of the ductile crust.

We conclude that emplacement of the HHG has been i) enhanced by the high magma buoyancy and ii) restricted to a brittlely deforming crust.

#### Rate of magma production

The volume of calc-alkaline magmas in the THB may be estimated from outcrop areas with an average depth of 7 km, as for calcalkaline batholiths (Haederle & Atherton, 2001). Calculation provides a magma production of about 1000 km<sup>3</sup>/km. Given the 14 Myr period of major production (Pettersson & Windley, 1985; Schärer et al., 1990), the rate of magma production was about 30 km<sup>3</sup>/km/Myr (Fig. 1). It plots within the low field of the CAB, which shows that calculation is fair.

We already estimated HHG volumes. Time is restricted to the interval 24-17 Ma (Scaillet et al., 1995a; Inger, 1994; Edwards & Harrison, 1997; Searle et al., 1997). The

rate production for the HHG is less than  $10 \text{ km}^3/\text{km}/\text{Ma}$ , about one order of magnitude less than island arcs (Fig. 1).

## Hercynian granites

The 360 to 285 Ma old Hercynian granites outcrop over about 3000 km from Spain to Bohemia, a length as long as the Himalayas. We adopt the division in Older (OIC) and Younger Intrusive Complexes (YIC) to distinguish two generations of Carboniferous granites separated by a time gap (Breiter et al., 1995; Siebel et al., 1997; Vignerresse, 2001). In Bohemia, the OIC are 325-315 Ma old and the YIC 305-285 Ma. In western France, granite ages span over a longer time (Peucat, 1982; Vidal et al., 1984). The youngest 310-285 Ma granites are located in southern Brittany only. The older granites span 360-315 Ma and are located in central Brittany. In the French Massif Central, the 335-325 Ma granites include most plutons in the centre of the Massif Central, whereas the 325-285 Ma ones outcrop all over the region (Duthou et al., 1984; Ploquin et al., 1994). In Spain, most granites intruded between 330-310 Ma. However, a 305-285 Ma magmatic event is also identified (Bellido Mulas et al., 1987).

The granodioritic OIC progressively grade into the YIC leucogranites, first with biotite dominant and then with two-mica granites, the muscovite content increasing with magma evolution. The OIC and YIC thus reflect two distinct causes for granite generation.

## Shapes

Hercynian plutons present 40 - 30 km lateral extensions. They are intrusive into Late Precambrian to Paleozoic rocks. Metamorphic aureoles generally indicate intrusion within the shallow brittle crust (Vignerresse, 1999).

The pluton shapes have been constrained from gravity data (Fig. 3). Three types reflect three levels of intrusion and deformation style environments.

- Type 1 plutons present an about 5 km deep floor gently dipping toward a root zone, marked by a sudden change in the slope of the floor, which reflects the brittle/ductile transition (Améglio et al., 1997; Fig. 3). The magma conduits were in the ductile crust, perpendicular to the major stress component in the brittle crust. In the brittle crust, the magma inflated against  $\sigma_3$  and plutons are ellipsoidal with long axes orthogonal to the maximum principal stress.
- Type 2 plutons have steep walls constrained by adjacent shear zones that indicate that magma intruded in a transtensive environment (e.g. Guineberteau et al., 1987). Most of these plutons have flat floors as deep as 10 km, with no evident root zone.
- Type 3 plutons are very thin and present a root zone parallel to the major stress component  $\sigma_1$  (e.g. Aranguren et al., 1986). They were emplaced within the brittle crust.

## Thermal conditions

Thermal conditions ruling Hercynian granite generation are constrained from the source, the depth of the ductile-brittle transition at time of emplacement, and the metamorphic aureoles.

High values of initial Sr, and the scatter of initial  $\epsilon_{\text{Nd}}$  attest that the source material is essentially crustal (Peucat, 1982; Vidal et al., 1984; Bernard Griffiths et al., 1985; Siebel et al., 1997; Jahn et al., 2000). The present-day interpretation invokes breakdown of hydrous minerals (Thompson, 1982; Vielzeuf & Holloway, 1988; Förster et al., 1999). The degree of melting of the source is high (above 30%), since it involved biotite breakdown. Accordingly, magma is estimated to have formed at  $850^\circ\text{C}$  and 600-700 MPa, corresponding to a depth of 22-25 km. The restitic level would correspond to the granulitic layer identified today at 18 km depth (e.g. Vignerresse, 1990).

The brittle/ductile transition occurs between 300-350  $^\circ\text{C}$  for quartz and above  $450^\circ\text{C}$  for feldspars (Kirby & Kronenberg, 1987). Type 1 shapes suggest that this paleo-transition is now 4 to 6 km deep. Metamorphic aureoles on the pluton tops are marked by illite formation at about  $250^\circ\text{C}$  and 200 MPa (Vignerresse, 1999). The thermal gradient acting at the time of granite emplacement can therefore be constrained to  $35^\circ\text{C}/\text{km}$ , in particular in Brittany (Vignerresse, 1999).

## Mechanical conditions

Hercynian plutons were emplaced within the brittle crust (Vignerresse et al., 1999). Both, shape and stress patterns may be estimated from the regional deformation field and from magmatic structures. In the brittle crust, the opening plane is perpendicular to the least principal stress  $\sigma_3$ . Because the vertical load is often the dominant stress component, the opening plane is vertical (Parsons et al., 1992). Magma raising along vertical planes starts crystallising. The loose framework formed by solid crystals allows stress transmission while magma is able to resist to the external stress field (Vignerresse et al., 1996). It can also slightly modify the near-field by increasing  $\sigma_3$ , and the intermediate principal stress  $\sigma_2$  (Vignerresse et al., 1999). If the magnitude of  $\sigma_3$  overcomes  $\sigma_2$ , magma intruding a vertical plane may locally open and pervade a horizontal plane (Parsons et al., 1992; Hogan et al., 1998).

## Rate of magma production

The production rate has been computed from outcrops, given an average depth of 5 km and spacing between plutons of about 70 km (Vignerresse, 1995). It provides a value of  $600 \text{ km}^3/\text{km}$  during about 20 Ma (Fig. 2). The corresponding rate of  $30 \text{ km}^3/\text{km}/\text{Ma}$  plots at the low end of magmatic arcs (Fig. 2).

## Discussion

Hercynian collision granites have been interpreted to be closest to Himalayan leucogranites (Sylvester, 1998). However, the HHG show a very low production rate. We now examine in more details their generation conditions.

### Source material and melting

#### Himalayan granites

Several models have been suggested to explain the HHG whose material source would be Formation I (Hodges et al., 1988; Inger & Harris, 1993; Rai et al., 1998).

Early interpretations related melting to the release of fluids from the footwall sequences during thrusting along the MCT (Le Fort, 1981; Le Fort et al., 1987; Harris, 1993). H<sub>2</sub>O and CO<sub>2</sub> produced during prograde metamorphism of footwall rocks built up in pressure, up to the lithostatic load. They could escape across the MCT and induce partial melting of the overlying hotter hanging wall because fluids decrease solidus temperatures (Le Fort, 1981). In this model, melt initiates as soon as the about 630°C solidus temperature is reached. The volume of leucogranitic magma requires melting of 10 to 15 % of the HHC (Le Fort, 1981). The interpretation is supported by the chemical composition in major and trace elements, and Sr, Nd and Pb variations in the HHG (Le Fort, 1981; Vidal et al., 1982; Le Fort et al., 1987; France Lanord & Le Fort, 1988).

Later interpretations questioned the abundant release of fluids. Indeed, evidence for pervasive aqueous fluid is scarce. Non-conventional values of incompatible elements (Rb, Sr and Ba) also make unlikely the fluid influence in HHG generation (Harris et al., 1993) and rare earth element patterns argue against fluid-saturated melting (Harris & Inger, 1992). Melting features are seen in the upper levels of the HHC, arguing against anatexis due to upward fluid influx, which would have resulted in melting close to the bottom MCT. Dehydration melting is more satisfying. The required water is released first from muscovite, then from biotite breakdowns and is immediately reincorporated into the magma (Thompson, 1982; Le Breton & Thompson, 1988). Experiments providing melt with the required composition have validated the process (Patiño Douce & Beard, 1995; Patiño & Harris, 1998). Formation I as the source has the adequate composition. Neodymium values ( $\epsilon_{Nd}$ ) range -11 to -18 reflecting the absence of mantle-derived magmas (Vidal et al., 1984; Inger & Harris, 1993; Ayres et al., 1997). They compare with  $\epsilon_{Nd}$  values measured on the HHC (-13 to -19). Rare earth studies corroborate partial and inhomogeneous melting of the source. Monazite presents incompletely re-homogenised core, which explains difficulties in obtaining reproducible Rb/Sr ages and initial  $I_{Sr}$  on the HHG (Vidal et al., 1984).

Equations 1 to 4 provide constraints on Formation I generating magma. The modal composition of Formation I indicates the presence of quartz (40%), plagioclase (11%), muscovite (22%), biotite (13%) and about 14 % garnet with a few accessory minerals (Ayres et al., 1997). The low proportion (c. 3%) of biotite in HHG is satisfactorily

explained if muscovite dehydration only operates. The temperature reached by the top Formation I is thus constrained at 750 °C. Owing to the low fluid phase coming from below the MCT, the reaction is buffered by the muscovite, quartz and plagioclase content. The total amount of melt that could be produced if the reaction was complete is about 2/3 of the initial source volume. Since minerals are not distributed in stoichiometric proportion, fertility reduces accordingly (Thompson, 1996; Vigneresse et al., 2001). Using the coefficients of Eq. 3, the low abundance of plagioclase and muscovite in Formation I constrains the potential volume of magma to about 13 %. Again, this value assumes complete consumption of these minerals, which was not the case. Nevertheless, the volume represented by 13 % of melting from the 10 km thick Formation I represents an about 1.3 km thick magma "layer", which matches the HHG thickness if they are in proportion of 70 % in the Formation III. This proportion compared to outcrop sizes is too high by a factor of 2. Therefore, the total melting proportion of Formation I should be less than 10 %.

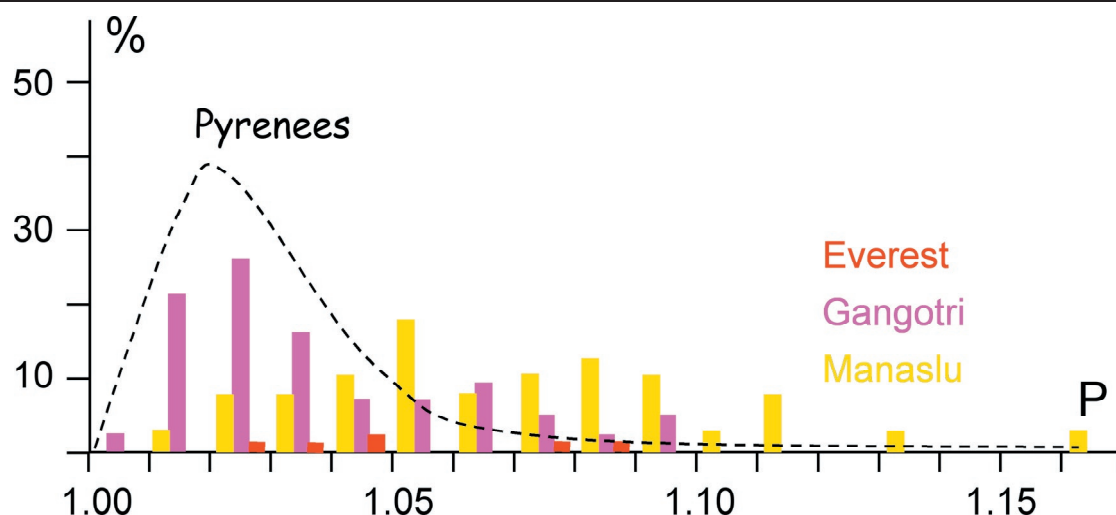
#### European Hercynian granites

Mafic enclaves are abundant in the OIC and rare in the YIC (Schödlbauer et al., 1997). These enclaves suggest that mantle had a significant role in generating the Hercynian granites.

Xenoliths scavenged by Tertiary volcanism in the French Massif Central (Maury et al., 1978; Downes et al., 1990) and in Saxony (Franz et al., 1997; Franz & Seifert, 1998) include granulite facies gneisses whose U-Pb ages correspond to the ages of the Hercynian granites. Therefore, Hercynian granites were produced by breakdown reactions leading to restitic granulites. The ubiquitous presence of biotite further indicates that source melting involved biotite breakdown around 850°C. The source material has not been precisely identified. Experimental melting of Precambrian protolith produces melt with the required composition (Vielzeuf & Montel, 1994; Castro et al., 2000). Such melt could reach 40 % at 300 MPa and 900°C, but it reduces to 15% if pressure >1500 MPa (Castro et al., 2000). In addition, temperatures >875°C are unlikely since they would correspond to garnet formation, which is only rarely observed.

Provided the thermal gradient of 35 °C/km (Fig. 2), the isotherm 850°C was 24.3 km deep, corresponding to 680 MPa. About 15-20 % of melt could be produced under such conditions.

Melting was in the same range as the average mineral proportion in sediments of the Hercynian crust (Vigneresse et al., 2001). Precambrian meta-sediments in Brittany (Denis & Dabard, 1988; Dabard et al., 1996) have high silica (70-80 %) and poor (maximum 7 %) plagioclase contents. They are a low fertile source (15-20 % melt at maximum). Precambrian gneiss in Saxony have less quartz (20-30 %) and more plagioclase (30-40 %; Tichomirowa et al., 2001). However the small amount of micas (7-20 % biotite and < 10% muscovite) does not preclude a high crustal fertility. In addition, the presence



**Figure 5.** ASM histogram in the Everest (black), Gangotri (grey) and Manaslu (pale grey) massifs. A global curve of anisotropy (P values) is also shown for Hercynian Pyrenean granites.

of about 8-15 % K-feldspar, an inert phase in reactions 2 and 4, diminishes fertility of the source region.

### Emplacement

#### Himalayan granites

The HHG emplaced in the ductile crust within a crustal segment undergoing southward extrusion (Hodges et al., 1993). A formulation that rules granite interaction with the local stress pattern, as under brittle conditions (Vignerresse et al., 1999) cannot apply. This point rules out the connection that has been suggested between magma emplacement and a crustal-scale gash (Guillot et al., 1999).

ASM measurements in the Everest (Rochette et al., 1994), Manaslu (Guillot et al., 1993) and Gangotri (Scaillet et al., 1995a) leucogranites reflect the emplacement anisotropy. The ellipsoid of magnetic susceptibility (axes  $K_1 > K_2 > K_3$ ) mimics the late magmatic strain ellipsoid. Anisotropy ( $P = K_1/K_3$ ) values range from 1.0 to 1.17 (Fig. 5). High P values document highest strain at the top of the Gangotri pluton (Scaillet et al., 1995a), which is consistent with the CS structures indicating that plutonism was synchronous with the STDS normal faulting (Burg et al., 1984; Burchfield et al., 1992). By comparison, P values of Hercynian leucogranites are generally lower than 6 % (Gleizes et al., 1997; Saint Blanquat et al., 2001).

Melt is channeled into foliation planes of the bottom Formation III, tapped and collected in small lenticular plutons against carbonates of the top Formation III. Carbonates have liberated  $\text{CO}_2$  during heating by granitic magmas, which drastically reduced the magma liquidus temperature (Holtz & Johannes, 1994). Thus magma crystallised and became rigid enough to stop moving and record solid-state CS structures.

The HHG represent aggregation of magma pulses reflected in the large variability of initial Sr ratios from sample to sample (Deniel et al., 1987). Pulses either came from different source regions, or from a source that experienced a continuously evolving degree of partial melting. Rare earth elements in zircons and monazites show the high melting variability of those minerals (Brouand et al.,

1990; Barbey et al., 1995), an argument for continuous melt extraction and progressive segregation into larger melt pockets (Barbey et al., 1996).

Aggregation is a consequence of the large HHC deformation and constrains tectonic models.

Successive loading and unloading on out of sequence thrusting in the extrusion model (Hodges et al., 1996) suffers from the very high load required to induce lateral and upward movements in the HHC (Hodges et al., 1996; Grujic et al., 1996). No décollement that would help initiating thrust motion was observed in the field. Granitic bodies, which can be considered as large strain markers (Gleizes et al., 1997), do not reflect the orientation of the major driving force: they are concordant to the foliation planes and thus oblique to the vertical load.

The numerical model driven by a very weak and overheated HHC lithology, with high heat production down to the 70 km deep Moho (Beaumont et al., 2001) implies boundary conditions that amplify the role of the weak layer during extrusion. The small amount of melt (< 13 %) cannot explain large weakening of the HHC.

The upper crust detached from the mantle at the upper/lower crust interface implies that buoyancy forces are the major driving forces. (Chemenda et al., 1995; 2000). In this scheme, the degree of melting of the HHC always remains low, and is consequently able neither to drive the large-scale weakening of the HHC, nor to induce shear deformation on thrust planes.

In our model, magma production is not required for initiating either the MCT or the STDS. Shear deformation contributes to magma aggregation, but magma is too viscous to be homogenised. The bulk shape of the HHG, conform to the foliation within Formation III, oblique to the STDS. This structural relationship precludes that loading / unloading control, which would require subhorizontal plutons perpendicular to the gravity load.

#### European Hercynian granites

Hercynian granites were mostly emplaced close to the brittle/ductile transition, which constituted a barrier for internal magma forces (Améglio et al., 1997). Their



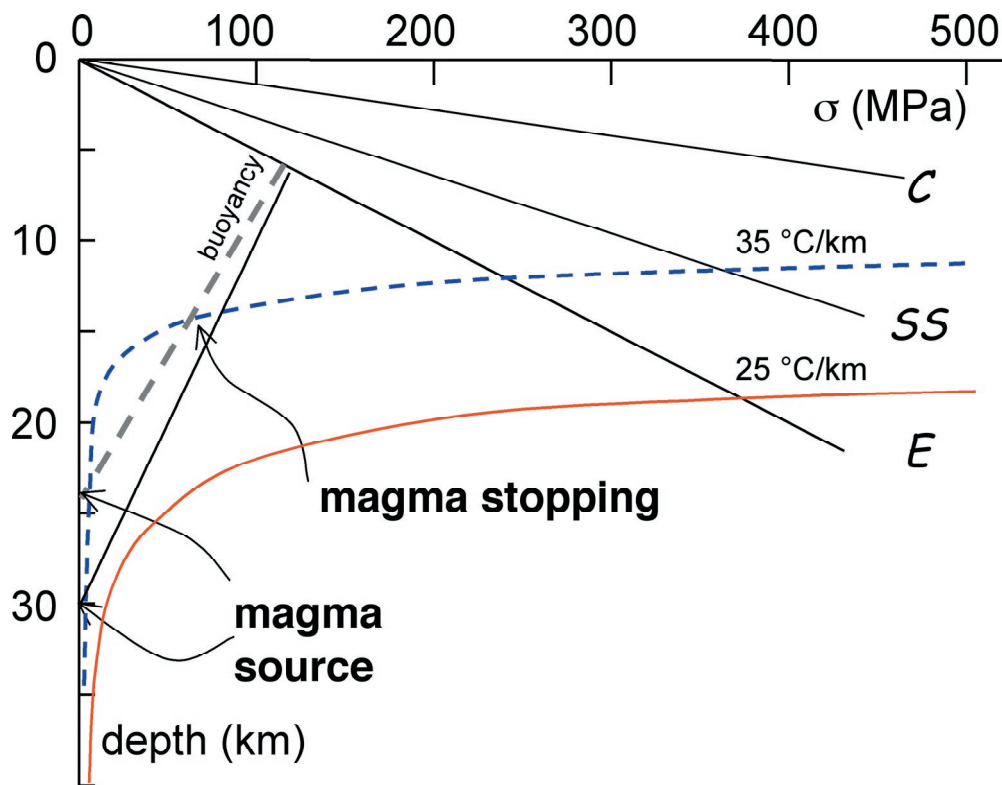


Fig. 6. Stress-Depth diagrams showing the various estimates of source and emplacement inferred for the Himalayan and Hercynian leucogranites according to the respective geothermal gradients (25 and 35 °C/km). . The extensional (E), strike slip (SS) and compressional (C) stress values for the brittle field are indicated. Temperature and pressure brackets are indicated from the contact metamorphism. The source region for magma generation is also reported. The black full lines correspond to a cold crust, the Himalayas. In contrast, the Hercynian conditions are in grey, with a lower strain rate and slightly increased density contrast between melt and its matrix. However, the important point in this diagram is the respective position of the rheological curve and the magmatic buoyancy curve.

emplacement was thus controlled by the local stress field (Vigneresse et al., 1999). Their internal fabrics, which records the last strain increment, may serve as a marker for the global strain pattern (Gleizes et al., 1997).

#### *Magma segregation and ascent*

We believe that the main differences between HHG and other collision-related granites relate to the conditions of their collect and ascent.

##### Himalayan granites

The Himalayan crust with its estimated geothermal gradient of 25°C/km was cold during granite generation. Melting temperature did not overcome 750°C at about 30 km below surface.

##### European Hercynian granites

In contrast, the Hercynian crust had a warm thermal gradient of 35 °C/km. Melting developed up to biotite breakdown, implying 850°C at about 24.3 km depth.

At a similar deformation rate, the brittle/ductile transition is shallower in the warmer crust, but it gets deeper with increasing strain rate (Fig. 2). A general rule is that an increase of about 2 orders of magnitude in strain rate (say from  $10^{-14}$  to  $10^{-12}$  s<sup>-1</sup>) is roughly equivalent to increasing the thermal gradient by 5 °C/km. The relationships between strain rate dependent viscosity and temperature

also interact with the hydrodynamic forces developed by a magma column, while remaining within the ductile field. In a cold crust, as the Himalayas, the magma column was high, but the viscous curve shows large stress values (Fig. 6). In consequence the hydrodynamic forces were always lower than stress required for viscous strain. Conversely, for a warm crust, as in the Hercynian, the viscous rheology was shallower, and due to the increased thermal gradient, its slope is steeper for low stress values. The buoyancy curve may intersect the curve of crustal rheology that will determine magma stopping. In the Hercynian crust, this took place at about 14.8 km (Fig. 6)

We also tested the change in rheology induced by the different rock types used in numerical models (Beaumont et al., 2001). Two end-members have been drawn, for a geothermal gradient of 20 and 40 °C/km (Fig. 2). High activation energy (485 kJ/mole) and the high power coefficient ( $n = 4.7$ ) introduced in numerical modeling (Beaumont et al., 2001) result in rheological characteristics that disagree with field observation. In a cold crust, the brittle/ductile transition is very deep (27-38 km). In the case of a hot crust (40 °C/km), the brittle/ductile transition is immediately above the melting zone at about 750 °C, which is not measured in rocks.

## Conclusions

The HHG are considerably smaller than the Hercynian and other collision-related granites. We recognize several reasons:

- 1) Melting was quite low (< 13 %), owing to the temperature below 780 °C, below the biotite breakdown. The volume of magma remained small because of this temperature, and also a poor fertile crustal and only crustal source.
- 2) The rate of magma production is considerably low.
- 3) Magma ascent has been controlled by the southward expulsion of the HHC.
- 4) The HHG result from the aggregation of successive pulses of magma of various origin, or different degree of melting. Magma re-homogenisation could not develop, leaving strong variation in isotopic values in near-by samples.
- 5) Magma has been able to reach the upper brittle crust. Emplacement was partly controlled by the contact with the calcareous black gneiss, from which CO<sub>2</sub> is released and lowers the melt below its solidus.

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